

## Sheet VII

Due: week of November 26

**Question 1** [*Energy-momentum-stress tensor for electromagnetic field*]:

(i) Derive the energy-momentum-stress tensor for the electromagnetic field

$$T_E^{\mu\nu} = \frac{1}{4\pi} \left( F^{\mu\alpha} F^\nu{}_\alpha - \frac{1}{4} (g^{-1})^{\mu\nu} F_{\alpha\beta} F^{\alpha\beta} \right), \quad (1)$$

where  $F_{\mu\nu}$  is the electromagnetic field strength tensor.

Hint: Use the definition

$$\frac{\partial}{\partial g_{\mu\nu}} (L_E d\mu_g) = -\frac{1}{2} T_E^{\mu\nu} d\mu_g, \quad (2)$$

where  $L_E$  is the electromagnetic Lagrangian

$$L_E = \frac{1}{16\pi} F^{\alpha\beta} F_{\alpha\beta} \quad (3)$$

and  $d\mu_g$  is the volume element

$$d\mu_g = \sqrt{-\det g} d^4x. \quad (4)$$

(ii) Let now

$$T^{\mu\nu} = T_M^{\mu\nu} + T_E^{\mu\nu}, \quad (5)$$

where  $T_M$  is the energy-momentum-stress tensor of matter. Use  $\nabla_\nu T^{\mu\nu} = 0$  and the definition of the spacetime force  $f^\mu$

$$f^\mu := \nabla_\nu T_M^{\mu\nu} \quad (6)$$

to show that

$$f^\mu = F^\mu{}_\nu J^\nu, \quad (7)$$

where  $J^\mu$  is the spacetime current. Now restrict to Minkowski spacetime. Find the time component  $f^0$  and space components  $f^i$  in terms of  $E^i$ ,  $B^i$ ,  $\rho$  and  $J^i$ , where  $E^i$  is the electric field,  $B^i$  is the magnetic field,  $J^i$  is the electric current and  $\rho$  is the charge density.

Hint: Use Maxwell's equations

$$\nabla_\nu F^{\mu\nu} = 4\pi J^\mu. \quad (8)$$

**Question 2** [*Geodesics in 3+1 formulation*]: The geodesic equation can be derived as critical points of the action

$$L = \int_a^b \sqrt{-g_{\mu\nu} \frac{dx^\mu}{d\lambda} \frac{dx^\nu}{d\lambda}} d\lambda. \quad (9)$$

Suppose we use a maximal time function  $t$  as the parameter for the geodesic, i.e.  $\lambda = t$ . Show that in the limit as  $c \rightarrow \infty$ , up to addition of a constant and multiplication with a constant, (9) becomes the action of a particle of unit mass in the gravitational field with potential  $\psi$ .

Hint: The metric is given by

$$g_{\mu\nu} dx^\mu dx^\nu = -c^2 \phi^2 dt^2 + \bar{g}_{ij} dx^i dx^j, \quad (10)$$

where

$$\phi = 1 + c^{-2} \psi + \dots, \quad \bar{g}_{ij} = \delta_{ij} + \dots \quad (11)$$

and  $\dots$  denotes higher order terms.

**Question 3** [*Kasner metric*]: We look at the Kasner metric

$$g = -dt^2 + \bar{g}, \quad (12)$$

where

$$\bar{g} = \sum_{i=1}^3 t^{2p_i} (dx^i)^2 \quad (13)$$

and  $t$  is the temporal distance function from a given spacelike hypersurface

$$(g^{-1})^{\mu\nu} \partial_\mu t \partial_\nu t = -1. \quad (14)$$

- (i) Calculate the components of the second fundamental form  $k_{ij}$ .
- (ii) Calculate the components of the Ricci tensor  $\bar{R}_{ij}$ .
- (iii) Show that (12) is a vacuum solution of the Einstein field equations if and only if

$$\sum_{i=1}^3 p_i = \sum_{i=1}^3 p_i^2 = 1. \quad (15)$$

Hint: Use the Einstein field equations in 3+1 form once we chose a foliation of the level sets of the temporal distance function.

**Question 4** [*Scalar curvature of maximal hypersurface*]: Suppose we have a maximal foliation. Look at the scalar curvature of a maximal hypersurface. Show that  $\bar{R} \geq 0$  with  $\bar{R} = 0$  at a point if and only if  $k_{ij} = 0$  and  $\rho = 0$  at this point.