

**Exercise 6.1 The Ising Paramagnet**

Consider  $N$  localized non-interacting magnetic moments which take the values  $s_i = \pm s$ . In the presence of an external magnetic field  $h$  the Hamiltonian of this system is given by

$$\mathcal{H} = - \sum_i h s_i. \quad (1)$$

Use the canonical ensemble to calculate the free energy  $F(T, h)$ , the caloric and thermal equations of state, the heat capacity  $C(T, h)$  and the magnetic susceptibility  $\chi(T, h)$ . Find a relation between the fluctuations in the magnetization and the susceptibility.

**Exercise 6.2 Non-interacting Particles in the Gravitational Field**

Consider a gas of non-interacting particles in the gravitational field

$$V_{\text{grav}}(x, y, z \geq 0) = mgz \quad (2)$$

with the gravitational constant  $g > 0$  at fixed temperature  $T$ . The volume of the gas is confined to a vertical, cylindrical vessel (radius  $R$ ) of semi-infinite height.

- a) Using the canonical ensemble, find the Helmholtz free energy, the entropy, and the internal energy of this system.
- b) Consider the system from the viewpoint of a local thermal equilibrium. Find the local particle density at height  $z$ ,  $n(z)$ , normalized such that  $N = \int d^3q n(q)$ . Find the local pressure  $p(z)$  as well as the local internal energy density  $u(z)$ . Express  $p(z)$  and  $u(z)$  in terms of  $n(z)$  to find the local caloric and thermal equations of state.
- c) Calculate the heat capacity, interpret and compare the different results using:
  - i) the entropy found in a).
  - ii) the equipartition law for  $U = \langle \mathcal{H} \rangle$ .
  - iii) the local caloric and thermal equations of state.
  - iv) the variance  $(\Delta \mathcal{H})^2$ .

*Hints for:*

- ii) Rewrite the Hamiltonian such that the equipartition law as given in chapter 3.4.2 in the lecture notes may be applied directly.
- iii) Keeping the specific volume  $v(z) = N/n(z)$  constant is equivalent to a constant local density and thus the local version of (1.17) in the lecture notes combined with (1.44) is given by

$$c_p = \left( \frac{\partial u}{\partial T} \right)_n + \left\{ \left( \frac{\partial U}{\partial V} \right)_T + p \right\} \alpha = \left( \frac{\partial u}{\partial T} \right)_n + T \left( \frac{\partial p}{\partial T} \right)_n \alpha. \quad (3)$$

First, show that even with the local viewpoint the thermal expansion coefficient  $\alpha$  is given by  $1/T$ , independent of  $z$ .

- iv) Use 3.4.3 in the lecture notes and identify the kinetic and potential contributions to the variance. Define the total potential energy through  $\langle z \rangle$ .

**Office hour:** Monday, October 29, 8-10 am (Sarah Etter, HIT K 12.2)