

Ex 1 (a) In $d = 4 - 2\epsilon$ dimensions the coupling parameter g_R acquires a mass dependence after renormalisation:

$$[g_R] = \epsilon.$$

$$\Rightarrow \text{Can write } g_R = \tilde{g}_R \mu^\epsilon$$

where $\mu = \text{renormalisation scale}$ & now $[\tilde{g}_R] = 0$.

By abuse of notation one just writes: $g_R \mu^\epsilon$.

Then from the definition of g_0 & Z_{iF} it follows that:

$$\underline{\underline{g_0 = Z_2^{-1} Z_A^{-\frac{1}{2}} Z_{iF} g_R \mu^\epsilon}}$$

(b) $Z_i = 1 + \delta Z_i$, $i = A, 2, iF$ where $\delta Z_i = O(g_R^2)$

$$\begin{aligned} \text{then: } g_0 &= (1 + \delta Z_2)^{-1} (1 + \delta Z_A)^{-\frac{1}{2}} (1 + \delta Z_{iF}) g_R \mu^\epsilon \\ &= (1 - \delta Z_2 + O(g_R^2)) (1 - \frac{1}{2} \delta Z_A + O(g_R^2)) \\ &\quad \times (1 + \delta Z_{iF}) g_R \mu^\epsilon \\ &= \underline{\underline{(1 - \delta Z_2 - \frac{1}{2} \delta Z_A + \delta Z_{iF}) g_R \mu^\epsilon}} \end{aligned}$$

From sheet 10:
$$\begin{cases} \delta Z_A = -\frac{g_R^2}{16\pi^2} \frac{2}{3} \left(\frac{2}{3} n_f - \frac{5}{3} N_c \right) \left(\frac{1}{\epsilon} + \text{finite} \right) \\ \delta Z_2 = -\frac{g_R^2}{16\pi^2} C_F \left(\frac{1}{\epsilon} + \text{finite} \right) \end{cases}$$

From sheet 11:
$$\begin{aligned} \delta Z_{iF} &= -\frac{g_R^2}{16\pi^2} (N_c + C_F) (4\pi)^{\epsilon} \Gamma(\epsilon) \\ &\simeq \underline{\underline{-\frac{g_R^2}{16\pi^2} (N_c + C_F) \left(\frac{1}{\epsilon} + \text{finite} \right)}} \quad (\text{in } \mathcal{J}=1 \text{ gauge}). \end{aligned}$$

$$\begin{aligned} \Rightarrow g_0 &= \left(1 + \frac{g_R^2}{16\pi^2} C_F \cdot \frac{1}{\epsilon} + \frac{1}{2} \cdot \frac{g_R^2}{16\pi^2} \left(\frac{2}{3} n_f - \frac{5}{3} N_c \right) \cdot \frac{1}{\epsilon} - \frac{g_R^2}{16\pi^2} (N_c + C_F) \frac{1}{\epsilon} + \text{finite} \right) g_R \mu^\epsilon \\ &= \underline{\underline{\left[1 - \frac{g_R^2}{16\pi^2} \left(\frac{11}{6} N_c - \frac{1}{3} n_f \right) \left(\frac{1}{\epsilon} + \text{finite} \right) \right] g_R \mu^\epsilon}} \end{aligned}$$

Exercise 2

(a) Using the result for g_0 in Ex 1 one has:

$$\begin{aligned} \mu \frac{\partial}{\partial \mu} \Big|_{g_0 = \text{const}} g_0 &= 0 \stackrel{!}{=} \epsilon \mu^\epsilon \left[1 - \frac{g_0^2}{16\pi^2} \left(\frac{11}{6} N_c - \frac{1}{3} n_f \right) \left(\frac{1}{\epsilon} + \text{finite} \right) \right] g_0 \\ &\quad + \mu^\epsilon \left[1 - \frac{3g_0^2}{16\pi^2} \left(\frac{11}{6} N_c - \frac{1}{3} n_f \right) \left(\frac{1}{\epsilon} + \text{finite} \right) \right] \beta \end{aligned}$$

$$\begin{aligned} \Rightarrow \beta &= -\epsilon g_0 \left[1 - \frac{g_0^2}{16\pi^2} \left(\frac{11}{6} N_c - \frac{1}{3} n_f \right) \left(\frac{1}{\epsilon} + \text{finite} \right) \right] \\ &\quad \times \left[1 - \frac{3g_0^2}{16\pi^2} \left(\frac{11}{6} N_c - \frac{1}{3} n_f \right) \left(\frac{1}{\epsilon} + \text{finite} \right) \right]^{-1} \\ &= -\epsilon g_0 \left[1 - \frac{g_0^2}{16\pi^2} \left(\frac{11}{6} N_c - \frac{1}{3} n_f \right) \left(\frac{1}{\epsilon} + \text{finite} \right) \right] \\ &\quad \times \left[1 + \frac{3g_0^2}{16\pi^2} \left(\frac{11}{6} N_c - \frac{1}{3} n_f \right) \left(\frac{1}{\epsilon} + \text{finite} \right) + \dots \right] \\ &\quad + O(g_0^5) \end{aligned}$$

$$\begin{aligned} \text{Let } \epsilon \rightarrow 0 \\ &= -2 \frac{g_0^3}{16\pi^2} \left(\frac{11}{6} N_c - \frac{1}{3} n_f \right) + O(g_0^5) \\ &= - \left(\frac{11}{3} N_c - \frac{2}{3} n_f \right) \frac{g_0^3}{16\pi^2} + O(g_0^5) \end{aligned}$$

(b) Let $b_0 = \frac{11}{3} N_c - \frac{2}{3} n_f$

then: $\mu \frac{\partial g_0}{\partial \mu} = -\frac{g_0^3}{16\pi^2} b_0 \Rightarrow g_0^2 = \frac{1}{\frac{b_0}{8\pi^2} \ln \mu + C}$, where $C = \text{const}$ of integration.

Using: $\lim_{\mu \rightarrow \Lambda_{\text{QCD}}} \frac{1}{g_0^2(\mu)} = 0 \Rightarrow C = -\frac{b_0}{8\pi^2} \ln \Lambda_{\text{QCD}}$.

$$\Rightarrow g_0^2 = \frac{1}{\frac{b_0}{8\pi^2} \ln \left(\frac{\mu}{\Lambda_{\text{QCD}}} \right)}$$

$$\Rightarrow \alpha_s = \frac{g_0^2}{4\pi} = \frac{4\pi}{\left(\frac{11}{3} N_c - \frac{2}{3} n_f \right) \ln \left(\frac{\mu}{\Lambda_{\text{QCD}}} \right)}$$

(c) In QCD $N_c=3, n_f=6 \Rightarrow \beta < 0 \Rightarrow$

Hence quarks behave like free particles at high energies!

