## Exercise 1. Free radiation field

The radiation field can be written as

$$\vec{A}(\vec{r},t) = \int \frac{d^3\vec{k}}{(2\pi)^3} \sum_{\lambda} \sqrt{\frac{2\pi\hbar c^2}{\omega_k}} \left( a_{k,\lambda}\vec{\varepsilon}_{k,\lambda} e^{i\vec{k}\cdot\vec{r}-i\omega_k t} + a_{k,\lambda}^*\vec{\varepsilon}_{k,\lambda}^* e^{-i\vec{k}\cdot\vec{r}+i\omega_k t} \right)$$

- 1. Compute the Hamiltonian  $H = \frac{1}{8\pi} \int d^3 \vec{r} (\vec{E}^2 + \vec{B}^2)$
- 2. Compute the Poynting vector  $\vec{P} = \frac{1}{4\pi c} \int d^3 \vec{r} \, \vec{E} \times \vec{B}$

## Exercise 2. Polarization vectors

Assuming the wave vector points along the z-axis,  $\vec{k} = (0, 0, k)$ , common choices for the polarization vectors  $\vec{\varepsilon}(\vec{k}, \lambda) = \vec{\varepsilon}_{k,\lambda}$  are: circular polarization:  $\vec{\varepsilon}_{k,+} = (-1, -i, 0)/\sqrt{2}$   $\vec{\varepsilon}_{k,-} = (1, -i, 0)/\sqrt{2}$ linear polarization:  $\vec{\varepsilon}_{k,1} = (1, 0, 0)$ ;  $\vec{\varepsilon}_{k,2} = (0, 1, 0)$ ;

1. Show

$$\vec{\varepsilon}^*(\vec{k},\lambda)\cdot\vec{\varepsilon}(\vec{k},\lambda')=\delta_{\lambda\lambda'}$$

2. Show

$$\sum_{\lambda} \varepsilon_i^*(\vec{k}, \lambda) \varepsilon_j(\vec{k}, \lambda) = \delta_{ij} - \frac{k_i k_j}{k^2}$$

## Exercise 3. Commutation relations

The Lagrangian density of the free radiation field is given by  $\mathcal{L} = -\frac{1}{16\pi}F^{\mu\nu}F_{\mu\nu}$  where  $F^{\mu\nu} \equiv \partial^{\mu}A^{\nu} - \partial^{\nu}A^{\mu}$  is the field strength tensor.

- 1. Show that the canonical momentum to  $\vec{A}$  is given by  $\vec{\pi} = -\frac{1}{4\pi c}\vec{E}$
- 2. After quantisation of the radiation field, both  $\vec{A}$  and  $\vec{\pi}$  are operators. Show that the commutation relations

$$\begin{aligned} &[\hat{a}(\vec{k},\lambda),\hat{a}^{\dagger}(\vec{k}',\lambda')] = (2\pi)^{3}\delta(\vec{k}-\vec{k}')\,\delta_{\lambda\lambda'} \\ &[\hat{a}(\vec{k},\lambda),\hat{a}(\vec{k}',\lambda')] = [\hat{a}^{\dagger}(\vec{k},\lambda),\hat{a}^{\dagger}(\vec{k}',\lambda')] = 0 \end{aligned}$$

result in equal-time commutation relations

$$[\hat{A}_{i}(\vec{x},t),\hat{\pi}_{j}(\vec{y},t)] = i\hbar \int \frac{d^{3}\vec{k}}{(2\pi)^{3}} e^{i\vec{k}\cdot(\vec{x}-\vec{y})} \left(\delta_{ij} - \frac{k_{i}k_{j}}{\vec{k}^{2}}\right)$$

How would these relations simplify if there were three polarizations of the radiation field? Interpret the results.