

Hands-On Session: Kelvin-Helmholtz Instability

In this exercise, we use the numerical code Ramses to explore the onset and evolution of the Kelvin-Helmholtz instability in a flow. To get up and running with the code, follow the instructions at <http://cheleb.net/dokuwiki/ramses>.

- (a) Look at the velocity field (v_x, v_y) in the first output. We see that v_y in the two zones differ (as expected), but we also see that $v_x \neq 0$, as we may have expected. Why is it not?
- (b) Look at the videos you have generated. What are the main features the Kelvin-Helmholtz instability evokes in the flow?
- (c) Recall the vorticity

$$\vec{\omega} = \nabla \times \vec{v}. \quad (1)$$

Derive the expression for the 2D vorticity ω , where $\vec{v} = (v_x, v_y, v_z) = (v_x, v_y, 0)$.

- (d) Recall from the lecture that we can rewrite the 2D vorticity equation as

$$\frac{D\omega}{Dt} = \nu \nabla^2 \omega. \quad (2)$$

Given that (for our particular simulation) Ramses is solving the Euler equations (i.e., there is no explicit viscosity ν specified anywhere in the code), how would you expect ω to evolve?

- (e) Take a look at the videos of ω and the y-averaged vorticity $\langle \omega \rangle_y$ you have created. How does $\langle \omega \rangle_y$ evolve? Does this contrast with your expectations? What does this imply in terms of ν ? How come?