Particle Physics Phenomenology II

FS 11, Series 9

Due date: 25.04.2011, 1 pm

Exercise 1 Consider neutrinos traveling through a medium of constant matter density ρ .

- i) Write down the Feynman diagrams of electron neutrinos scattering off electrons. Which diagram is responsible for neutrino oscillations?
- ii) The matter oscillation may be described by adding an additional term to the Hamiltonian in the flavor basis, such that

$$\mathcal{H}_F = U \operatorname{diag}(0, \Delta m_{21}^2/2E, \Delta m_{31}^2/2E) U^{\dagger} + \operatorname{diag}(V_{CC}, 0, 0)$$

where $V_{CC} = \pm \sqrt{2} G_F N_e$ (plus for neutrinos, minus for antineutrinos) is the *matter* potential with $N_e \simeq Y_e \rho/m_N$ the electron density in matter, Y_e the number of electrons per nucleon with mass m_N . In the two flavor case show that this Hamiltonian leads to an effective mixing angle satisfying

$$\tan 2\theta_m = \frac{\sin 2\theta_{12}}{\cos 2\theta_{12} - \frac{2EV_{CC}}{\Delta m_{21}^2}}$$

Hint: Ignore third components. Multiply out \mathcal{H}_F and subtract a suitable diagonal term, (why can you always do this?) which brings \mathcal{H}_F into a form such that $\mathcal{H}_{11} = -\mathcal{H}_{22}$ as in the vacuum case. Read of the modified $\cos 2\theta$.

- iii) Compute the Energy of the neutrino which leads to maximal mixing in the sun, where $\theta_{12} = 33.9^{\circ}$, $\Delta m_{21}^2 = 8 \cdot 10^5 \text{eV}^2$ and $\rho = 150 \text{gcm}^{-3}$.
- iv) Is the sign of Δm_{21}^2 important?
- v) The parameters Δm_{21}^2 and θ_{12} extracted from solar oscillations can also be obtained in laboratory experiments, namely accelerator and reactor experiments (with typical energies: 1 GeV for accelerator ν 's and 4 MeV for reactor ν 's). What is the optimal distance between neutrino source and detector for these two types of experiments to confirm the solar parameters?